# Fundamentals of Fluid Dynamics: Waves in Fluids

Introductory Course on Multiphysics Modelling

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(after: D.J. ACHESON's "Elementary Fluid Dynamics")

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- 1 Introduction
  - The notion of wave
  - Basic wave phenomena
  - Mathematical description of a traveling wave

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### The notion of wave

#### What is a wave?

A wave is the transport of a disturbance (or energy, or piece of information) in space not associated with motion of the medium occupying this space as a whole. (Except that electromagnetic waves require no medium !!!)

- The transport is at **finite speed**.
- The shape or form of the **disturbance** is **arbitrary**.
- The disturbance moves with respect to the medium.

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- Iongitudinal waves the disturbance moves parallel to the direction of propagation. *Examples*: sound waves, compressional elastic waves (P-waves in geophysics);
- transverse waves the disturbance moves perpendicular to the direction of propagation. Examples: waves on a string or membrane, shear waves (S-waves in geophysics), water waves, electromagnetic waves.

### Basic wave phenomena

```
    reflection – change of wave direction from hitting a reflective surface,
    refraction – change of wave direction from entering a new medium,
    diffraction – wave circular spreading from entering a small hole (of the wavelength-comparable size), or wave bending around small obstacles,
    interference – superposition of two waves that come into contact with each other.
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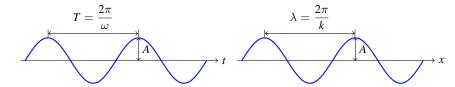
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### Standing wave

A **standing wave**, also known as a **stationary wave**, is a wave that remains in a constant position. This phenomenon can occur:

- when the medium is moving in the opposite direction to the wave,
- (in a stationary medium:) as a result of interference between two waves travelling in opposite directions.

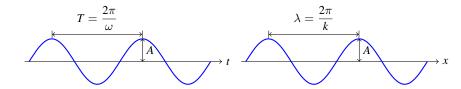


#### **Traveling waves**

Simple wave or **traveling wave**, sometimes also called *progressive* wave, is a disturbance that varies both with time t and distance x in the following way:

$$u(x,t) = A(x,t) \cos (kx - \omega t + \theta_0)$$
  
=  $A(x,t) \sin (kx - \omega t + \underbrace{\theta_0 \pm \frac{\pi}{2}}_{\widehat{\theta}_0})$ 

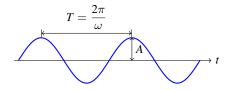
where A is the **amplitude**,  $\omega$  and k denote the **angular frequency** and **wavenumber**, and  $\theta_0$  (or  $\tilde{\theta}_0$ ) is the initial **phase**.



#### **Traveling waves**

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- **Amplitude** *A* [e.g. m, Pa, V/m] a measure of the maximum disturbance in the medium during one wave cycle (the maximum distance from the highest point of the crest to the equilibrium).
- **Phase**  $\theta = kx \omega t + \theta_0$  [rad], where  $\theta_0$  is the *initial* phase (shift), often ambiguously, called the phase.



- Period T [s] the time for one complete cycle for an oscillation of a wave.
- **Frequency** f [Hz] the number of periods per unit time.

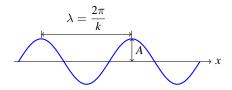
#### Frequency and angular frequency

The **frequency** f [Hz] represents the number of periods per unit time

$$f = \frac{1}{T} \,.$$

The **angular frequency**  $\omega$  [Hz] represents the frequency in terms of radians per second. It is related to the frequency by

$$\omega = \frac{2\pi}{T} = 2\pi f.$$



■ Wavelength  $\lambda$  [m] – the distance between two sequential crests (or troughs).

### Wavenumber and angular wavenumber

The **wavenumber** is the spatial analogue of frequency, that is, it is the measurement of the number of repeating units of a propagating wave (the number of times a wave has the same phase) per unit of space.

Application of a Fourier transformation on data as a function of time yields a frequency spectrum; application on data as a function of position yields a wavenumber spectrum.

The **angular wavenumber** k  $\left[\frac{1}{\mathrm{m}}\right]$ , often misleadingly abbreviated as "wave-number", is defined as  $k=\frac{2\pi}{\lambda}$ .

There are two velocities that are associated with waves:

**1 Phase velocity** – the rate at which the wave propagates:

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2 Group velocity – the velocity at which variations in the shape of the wave's amplitude (known as the modulation or envelope of the wave) propagate through space:

$$c_{\mathsf{g}} = rac{\mathrm{d}\omega}{\mathrm{d}k}$$
.

This is (in most cases) the signal velocity of the waveform, that is, the **rate at which information or energy is transmitted** by the wave. However, if the wave is travelling through an absorptive medium, this does not always hold.

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- Consider **two-dimensional** water waves:  $\mathbf{u} = [u(x, y, t), v(x, y, t), 0]$ .
- Suppose that the flow is **irrotational**:  $\frac{\partial v}{\partial x} \frac{\partial u}{\partial y} = 0$ .
- Therefore, there exists a **velocity potential**  $\phi(x, y, t)$  so that

$$u = \frac{\partial \phi}{\partial x}$$
,  $v = \frac{\partial \phi}{\partial y}$ .

The fluid is **incompressible**, so by the virtue of the incompressibility condition,  $\nabla \cdot \pmb{u} = 0$ , the velocity potential  $\phi$  will satisfy **Laplace's equation** 

$$\frac{\partial^2 \phi}{\partial x^2} + \frac{\partial^2 \phi}{\partial y^2} = 0.$$

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#### Free surface

The fluid motion arises from a deformation of the water surface. The equation of this free surface is denoted by  $y = \eta(x, t)$ .

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Kinematic condition at the free surface

#### Kinematic condition at the free surface:

Fluid particles on the surface must remain on the surface.

The kinematic condition entails that  $F(x,y,t) = y - \eta(x,t)$  **remains constant** (in fact, zero) for any particular particle on the free surface which means that

$$\frac{\mathrm{D}F}{\mathrm{D}t} = \frac{\partial F}{\partial t} + (\mathbf{u} \cdot \nabla)F = 0$$
 on  $y = \eta(x, t)$ ,

and this is equivalent to

$$\frac{\partial \eta}{\partial t} + u \frac{\partial \eta}{\partial x} = v \quad \text{on} \quad y = \eta(x, t).$$

Pressure condition at the free surface

#### Pressure condition at the free surface:

The fluid is **inviscid** (by assumption), so the condition at the free surface is simply that the pressure there is equal to the atmospheric pressure  $p_0$ :

$$p = p_0$$
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#### Bernoulli's equation for unsteady irrotational flow

If the flow is irrotational (so  $u=\nabla\phi$  and  $\nabla\times u=0$ ), then, by integrating (over the space domain) the **Euler's momentum equation**:

$$\frac{\partial \nabla \phi}{\partial t} = -\nabla \left( \frac{p}{\varrho} + \frac{1}{2} u^2 + \chi \right),\,$$

the Bernoulli's equation is obtained

$$\frac{\partial \phi}{\partial t} + \frac{p}{\rho} + \frac{1}{2} u^2 + \chi = G(t).$$

Here,  $\chi$  is the gravity potential (in the present context  $\chi=g\,y$  where g is the gravity acceleration) and G(t) is an arbitrary function of time alone (a constant of integration).

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$$\frac{\partial \phi}{\partial t} + \frac{p}{\rho} + \frac{1}{2} u^2 + g y = G(t).$$

Here, G(t) is an arbitrary function of time alone (a constant of integration).

Now, by choosing G(t) in a convenient manner,  $G(t)=\frac{p_0}{\varrho}$ , the **pressure condition** may be written as:  $\frac{\partial \phi}{\partial t} + \frac{1}{2} \left(u^2 + v^2\right) + g \, \eta = 0 \quad \text{on} \quad y = \eta(x,t).$ 

Small amplitude waves: the linearized surface conditions

#### **Small-amplitude waves**

The free surface displacement  $\eta(x,t)$  and the fluid velocities u,v are small.

Linearization of the kinematic condition

$$\begin{aligned} v &= \frac{\partial \eta}{\partial t} + \underbrace{u \frac{\partial \eta}{\partial x}}_{\text{small}} & \rightarrow & v(x, \eta, t) = \frac{\partial \eta}{\partial t} \\ & \xrightarrow{\text{Taylor}}_{\text{series}} & v(x, 0, t) + \underbrace{\eta \frac{\partial v}{\partial y} \left( x, 0, t \right) + \cdots}_{\text{small}} = \frac{\partial \eta}{\partial t} \\ & \rightarrow & v(x, 0, t) = \frac{\partial \eta}{\partial t} & \xrightarrow{\frac{v = \frac{\partial \phi}{\partial y}}{\partial y}} & \underbrace{\left( \frac{\partial \phi}{\partial y} = \frac{\partial \eta}{\partial t} \quad \text{on } y = 0. \right)}_{\text{small}} \end{aligned}$$

■ Linearization of the pressure condition

$$\frac{\partial \phi}{\partial t} + \underbrace{\frac{1}{2} (u^2 + v^2)}_{\text{small}} + g \eta = 0 \quad \rightarrow \quad \underbrace{\left( \frac{\partial \phi}{\partial t} + g \eta = 0 \quad \text{on } y = 0. \right)}_{\text{small}}$$

Dispersion relation and travelling wave solution

### A sinusoidal travelling wave solution

The free surface is of the form

$$\eta = A \cos(kx - \omega t),$$

where A is the **amplitude** of the surface displacement,  $\omega$  is the **circular frequency**, and k is the **circular wavenumber**.

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■ The corresponding velocity potential is

$$\phi = q(y) \sin(kx - \omega t)$$
.

- It satisfies the Laplace's equation,  $\frac{\partial^2 \phi}{\partial x^2} + \frac{\partial^2 \phi}{\partial x^2} = 0$ .
- $\blacksquare$  Therefore,  $q(\mathbf{y})$  must satisfy  $q''-k^2q=0$  , the general solution of which is

$$q = C \exp(ky) + D \exp(-ky).$$

For deep water waves D=0 (if k>0 which may be assumed without loss of generality) in order that the velocity be bounded as  $y\to -\infty$ .

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- Now, the (linearized) **free surface conditions** yield what follows:
  - 1 the kinematic condition  $(\frac{\partial \phi}{\partial y} = \frac{\partial \eta}{\partial t} \text{ on } y = 0)$ :

$$Ck = A\omega \rightarrow \left(\phi = \frac{A\omega}{k} \exp(ky) \sin(kx - \omega t)\right),$$

2 the pressure condition  $(\frac{\partial \phi}{\partial t} + g \eta = 0 \text{ on } y = 0)$ :

$$-C\omega + gA = 0 \rightarrow (\omega^2 = gk.)$$
 (dispersion relation!)

Particle paths

#### The fluid velocity components:

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Any particle departs only a **small amount** (X,Y) **from its mean position** (x,y). Therefore, its position as a function of time may be found by integrating  $u = \frac{dX}{dt}$  and  $v = \frac{dY}{dt}$ ; whence:

$$X(t) = -A \, \exp(k \, y) \, \sin(k \, x - \omega \, t) \,, \quad Y(t) = A \, \exp(k \, y) \, \cos(k \, x - \omega \, t) \,.$$

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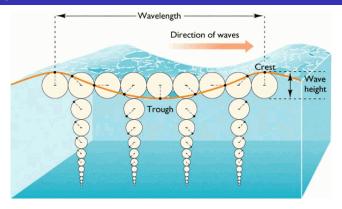
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- Particle paths are circular.
- The radius of the path circles,  $A \exp(ky)$ , decrease exponentially with depth. So do the fluid velocities.
- Virtually all the energy of a surface water wave is contained within half a wavelength below the surface.

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Effects of finite depth

#### Effects of finite depth

If the **fluid is bonded below** by a rigid plane y = -h, so that

$$v = \frac{\partial \phi}{\partial y} = 0$$
 at  $y = -h$ ,

the dispersion relation and the phase speed are as follows:

$$\omega^2 = g k \tanh(k h), \qquad c^2 = \frac{g}{k} \tanh(k h).$$

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There are two limit cases:

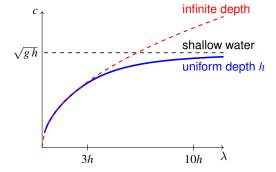
- 1  $h \gg \lambda$  (infinite depth):  $kh = 2\pi \frac{h}{\lambda}$  is large and  $\tanh(kh) \approx 1$ , so  $c^2 = \frac{g}{k}$ . In practice, this is a good approximation if  $h > \frac{1}{3}\lambda$ .
- 2  $h \ll \lambda/2\pi$  (shallow water):  $kh \ll 1$  and  $\tanh(kh) \approx kh$ , so  $c^2 = gh$ , which means that c is independent of k in this limit. Thus, the gravity waves in shallow water are non-dispersive.

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### Dispersion and the group velocity

### Dispersion of waves

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There are generally two sources of dispersion:

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- 2 the waveguide dispersion occurs when the speed of a wave in a waveguide depends on its frequency for geometric reasons, independent of any frequency-dependence of the materials from which it is constructed.

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### **Dispersion relation**

$$(\omega = \omega(k)) = c(k) k, \qquad c = c(k) = \frac{\omega(k)}{k}.$$

If  $\omega(k)$  is a **linear** function of k then c is **constant** and the medium is **non-dispersive**.

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• deep water waves:  $\omega = \sqrt{g \, k},$   $c = \sqrt{\frac{g}{k}}.$ 

▶ finite depth waves:  $\omega = \sqrt{g k \tanh(k h)}, \quad c = \sqrt{\frac{g}{k} \tanh(k h)}.$ 

▶ shallow water waves:  $\omega = \sqrt{g h} k$ ,  $c = \sqrt{g h} \rightarrow \text{non-dispersive!}$ 

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- ▶ finite depth waves:  $\omega = \sqrt{g \, k \, \mathrm{tanh}(k \, h)}, \quad c = \sqrt{\frac{g}{k} \, \mathrm{tanh}(k \, h)}.$
- lacktriangle shallow water waves:  $\omega = \sqrt{g\,h}\,k$ ,  $c = \sqrt{g\,h} \, o\,$  non-dispersive!

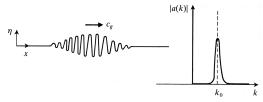
#### Group and phase velocity

$$c = \frac{\mathrm{d}\omega}{\mathrm{d}k}, \qquad c = \frac{\omega}{k}.$$

- In **dispersive systems** both velocities are different and frequency-dependent (i.e., wavenumber-dependent):  $c_q = c_q(k)$  and c = c(k).
- In **non-dispersive systems** they are equal and constant:  $c_q = c$ .

#### Important properties of the group velocity:

1 At this velocity the isolated wave packet travels as a whole.



**Discussion for a wave packet**: for k in the neighbourhood of  $k_0$ 

$$\omega(k) pprox \omega(k) + (k - k_0) \, c_{\mathsf{g}} \,, \quad \mathsf{where} \ \ c_{\mathsf{g}} = \left. \frac{\mathrm{d}\omega}{\mathrm{d}k} \, \right|_{k = k_0},$$

and  $\omega(k)=0$  outside the neighbourhood; the Fourier integral equals

$$\eta(x,t) = \operatorname{Re}\left[\int\limits_{-\infty}^{\infty} a(k) \, \exp\left(\mathrm{i} \, (k \, x - \omega \, t)\right) \, \mathrm{d}k\right] \quad \leftarrow \text{(for a general disturbance)}$$

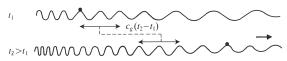
$$\approx \operatorname{Re}\left[\overbrace{\exp\left(\mathrm{i}\left(k_{0}\,x - \omega(k_{0})\,t\right)\right)}^{\operatorname{a pure harmonic wave}}\int_{0}^{\infty} \underbrace{a\,\mathrm{function of}\,(x - c_{\mathrm{g}}\,t)}_{a(k)\,\exp\left(\mathrm{i}\,(k - k_{0})\,(x - c_{\mathrm{g}}\,t)\right)}\mathrm{d}k\right].$$

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- 1 At this velocity the isolated wave packet travels as a whole.
- **2** The **energy is transported** at the group velocity.
- One must travel at the group velocity to see the waves of the same wavelength.



A slowly varying wavetrain can be written as

$$\eta(x,t) = \operatorname{Re}\left[A(x,t)\,\exp\left(\mathrm{i}\,\theta(x,t)\right)\right],$$

where the **phase function**  $\theta(x,t)$  describes the oscillatory aspect of the wave, while A(x,t) describes the gradual modulation of its amplitude.

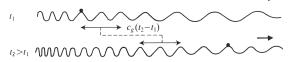
The local wavenumber and frequency are defined by

$$k = \frac{\partial \theta}{\partial x}$$
,  $\omega = -\frac{\partial \theta}{\partial t}$ .

For purely sinusoidal wave  $\theta = kx - \omega t$ , where k and  $\omega$  are constants.

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,  $\omega = -\frac{\partial \theta}{\partial t}$ .

For purely sinusoidal wave  $\theta = kx - \omega t$ , where k and  $\omega$  are constants. In general, k and  $\omega$  are functions of x and t. It follows immediately that

$$\frac{\partial k}{\partial t} + \frac{\partial \omega}{\partial x} = 0 \longrightarrow \frac{\partial k}{\partial t} + \frac{\mathrm{d}\omega}{\mathrm{d}k} \frac{\partial k}{\partial x} = \frac{\partial k}{\partial t} + c_{\mathsf{g}}(k) \frac{\partial k}{\partial x} = 0$$

which means that k(x, t) is constant for an observer moving with the velocity  $c_q(k)$ .

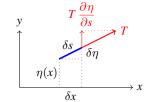
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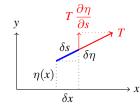
- The **vertical component** of surface tension force equals  $T \frac{\partial \eta}{\partial s}$ , where s denotes the distance along the surface.
- For small wave amplitudes  $\delta s \approx \delta x$ , and then  $T \frac{\partial \eta}{\partial s} \approx T \frac{\partial \eta}{\partial r}$ .



#### **Surface tension**

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■ A small portion of surface of length  $\delta x$  will experience surface tension at both ends, so the **net upward force** on it will be

$$T \left. \frac{\partial \eta}{\partial x} \right|_{x + \delta x} - T \left. \frac{\partial \eta}{\partial x} \right|_{x} = T \left. \frac{\partial^2 \eta}{\partial x^2} \, \delta x \right|_{x}$$

Therefore, an upward force **per unit area of surface** is  $T \frac{\partial^2 \eta}{\partial r^2}$ .

#### Local equilibrium at the free surface

The net upward force per unit area of surface,  $T \frac{\partial^2 \eta}{\partial x^2}$ , must be balanced by the difference between the atmospheric pressure  $p_0$  and the pressure p in the fluid just below the surface:

$$p_0 - p = T \frac{\partial^2 \eta}{\partial x^2}$$
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This **pressure condition** at the free surface takes into consideration the **effects of surface tension**. The kinematic condition remains the same: fluid particles cannot leave the surface.

#### Linearized free surface conditions (with surface tension effects)

For small amplitude waves:

$$\frac{\partial \phi}{\partial y} = \frac{\partial \eta}{\partial t}$$
,  $\frac{\partial \phi}{\partial t} + g \eta = \frac{T}{\rho} \frac{\partial^2 \eta}{\partial x^2}$  on  $y = 0$ .

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A sinusoidal travelling wave solution  $\eta = A \cos(kx - \omega t)$  leads now to a new **dispersion relation** 

$$\left(\omega^2 = g\,k + \frac{T\,k^3}{\varrho}.\right)$$

As a consequence, the **phase** and **group velocities** include now the surface tension effect:

$$c = rac{\omega}{k} = \sqrt{rac{g}{k} + rac{T\,k}{arrho}}\,, \qquad c_{\mathsf{g}} = rac{\mathrm{d}\omega}{\mathrm{d}k} = rac{g + 3T\,k^2/arrho}{2\sqrt{g\,k + T\,k^3/arrho}}\,.$$

#### Surface tension importance parameter

The relative importance of surface tension and gravitational forces in a fluid is measured by the following parameter

$$\beta = \frac{T \, k^2}{\varrho \, g} \, .$$

(The so-called *Bond number*  $= \frac{\varrho g L^2}{T}$ ; it equals  $\frac{4\pi^2}{\beta}$  if  $L = \lambda$ .)

Now, the dispersion relation, as well as the phase and group velocities can be written as

$$\omega^2 = g k (1 + \beta), \qquad c = \sqrt{\frac{g}{k} (1 + \beta)}, \qquad c_g = \frac{g (1 + 3\beta)}{2\sqrt{g k (1 + \beta)}}.$$

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Depending on the parameter  $\beta$ , two extreme cases are distinguished:

1  $\beta \ll$  1: the effects of surface tension are negligible – the waves are **gravity waves** for which

$$\omega^2 = g k$$
,  $c = \sqrt{\frac{g}{k}} = \sqrt{\frac{g \lambda}{2\pi}}$ ,  $c_g = \frac{c}{2}$ .

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 $\beta \gg 1$ : the waves are essentially **capillary waves** for which

$$\omega^2 = g\,k\,\beta = \frac{T\,k^3}{\varrho}, \quad c = \sqrt{\frac{g}{k}\beta} = \sqrt{\frac{T\,k}{\varrho}} = \sqrt{\frac{2\pi\,T}{\varrho\,\lambda}}\,, \quad c\mathrm{g} = \frac{g\,3\beta}{2\sqrt{g\,k\,\beta}} = \frac{3}{2}c\,.$$

**CAPILLARY WAVES:** 

short waves travel faster,

**GRAVITY WAVES:** 

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### CAPILLARY WAVES:

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- $\begin{tabular}{ll} \blacksquare & the group velocity exceeds the \\ & phase velocity, $c_{\rm g}>c$, \end{tabular}$

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### CAPILLARY WAVES:

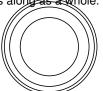
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wave patterns

#### **GRAVITY WAVES:**

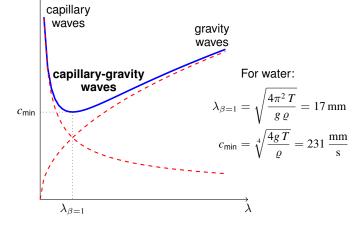
- long waves travel faster.
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The **capillary effects** predominate when **raindrops** fall on a pond, and as short waves travel faster the **wavelength decreases with radius** at any particular time.

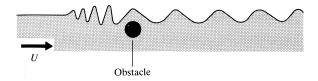
The effects of gravity predominate when a large stone is dropped into a pond, and as long waves travel faster the wavelength increases with radius at any particular time.

## Capillary-gravity waves



For  $\beta \approx 1$  both effects (the surface tension and gravity) are significant and the waves are **capillary-gravity waves**.

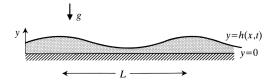
Example: Uniform flow past a submerged obstacle



- 1  $U < c_{\min}$  there are no steady waves generated by the obstacle;
- **2**  $U > c_{\min}$  there are **two values** of  $\lambda$  ( $\lambda_1 > \lambda_2$ ) for which c = U:  $\lambda_1$  **the larger value** represents a **gravity wave**:
  - $\blacksquare$  the corresponding group velocity is less than c,
  - the energy of this relatively long-wavelength disturbance is carried downstream of the obstacle.
  - $\lambda_2$  the smaller value represents a capillary wave:
    - $\blacksquare$  the corresponding group velocity is greater than c,
    - the energy of this relatively short-wavelength disturbance is carried upstream of the obstacle, where it is rather quickly dissipated by viscous effects, on account of the short wavelength (in fact, each wave-crest is at rest, but relative to still water it is travelling upstream with speed *U*).

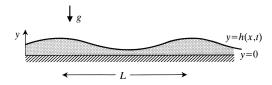
### Assumptions:

- The amplitudes of waves are finite, that is, not (infinitesimally) small compared with the depth; therefore, the linearized theory does not apply.
- A typical value  $h_0$  of depth h(x,t) is much smaller than a typical horizontal length scale L of the wave, that is:  $h_0 \ll L$ . This is the basis of the so-called **shallow-water approximation**.



► The full (nonlinear) 2-D equations are:

$$\frac{\mathrm{D}u}{\mathrm{D}t} = -\frac{1}{\varrho} \frac{\partial p}{\partial x} \;, \qquad \frac{\mathrm{D}v}{\mathrm{D}t} = -\frac{1}{\varrho} \frac{\partial p}{\partial y} - g \;, \qquad \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = 0 \;.$$



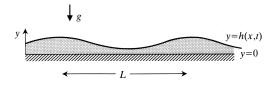
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▶ In the shallow-water approximation (when  $h_0 \ll L$ ) the vertical component of acceleration can be neglected in comparison with the gravitational acceleration:

$$\frac{\mathrm{D}v}{\mathrm{D}t} \ll g \quad \rightarrow \quad 0 = -\frac{1}{\varrho} \frac{\partial p}{\partial y} - g \quad \rightarrow \quad \frac{\partial p}{\partial y} = \varrho \, g \, .$$

Integrating and applying the condition  $p = p_0$  at y = h(x, t) gives

$$p(x, y, t) = p_0 - \varrho g [y - h(x, t)].$$



$$\frac{\mathrm{D}u}{\mathrm{D}t} = -\frac{1}{\varrho} \frac{\partial p}{\partial x} \;, \qquad \frac{\mathrm{D}v}{\mathrm{D}t} = -\frac{1}{\varrho} \frac{\partial p}{\partial y} - g \;, \qquad \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = 0 \;.$$

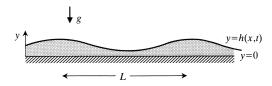
▶ In the shallow-water approximation (when  $h_0 \ll L$ ) the vertical component of acceleration can be neglected and then

$$p(x, y, t) = p_0 - \varrho g [y - h(x, t)].$$

This is used for the equation for the horizontal component of acceleration:

$$\frac{\mathrm{D}u}{\mathrm{D}t} = -g \frac{\partial h}{\partial x} \xrightarrow{\frac{\partial u}{\partial y} = 0} \left( \frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} = -g \frac{\partial h}{\partial x} \right)$$

where u = u(x, t) and h = h(x, t).



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➤ A second equation linking u and h may be obtained as follows:

$$\frac{\partial v}{\partial y} = -\,\frac{\partial u}{\partial x} \quad \to \quad v(x,y,t) = -\,\frac{\partial u(x,t)}{\partial x}\,\,y + f(x,t) \quad \xrightarrow{\nu = 0 \text{ at } y \,=\, 0} \quad v = -\,\frac{\partial u}{\partial x}\,\,y\,,$$

and using the **kinematic condition at the free surface** – fluid particles on the surface must remain on it, so the vertical component of velocity  $\nu$  equals the rate of change of the depth h when moving with the horizontal velocity u:

$$v = \frac{\partial h}{\partial t} + u \frac{\partial h}{\partial x}$$
 at  $y = h(x, t)$   $\rightarrow \left(\frac{\partial h}{\partial t} + u \frac{\partial h}{\partial x} + h \frac{\partial u}{\partial x} = 0\right)$ .

### Shallow-water equations

Nonlinear equations for the horizontal component of velocity u = u(x, t) and the depth h = h(x, t) of finite-amplitude waves on shallow water:

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + g \frac{\partial h}{\partial x} = 0, \qquad \frac{\partial h}{\partial t} + u \frac{\partial h}{\partial x} + h \frac{\partial u}{\partial x} = 0.$$

(The vertical component of velocity is  $v(x, y, t) = -\frac{\partial u}{\partial x} y$ .)

On introducing the new variable  $c(x,t) = \sqrt{g\,h}$  and then adding and subtracting the two equations the form suited to treatment by the *method of characteristics* is obtained

$$\left[\frac{\partial}{\partial t} + (u+c)\frac{\partial}{\partial x}\right](u+2c) = 0, \qquad \left[\frac{\partial}{\partial t} + (u-c)\frac{\partial}{\partial x}\right](u-2c) = 0.$$

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Let x = x(s), t = t(s) be a **characteristic curve** defined parametrically (s is the parameter) in the x-t plane and starting at some point  $(x_0, t_0)$ . In fact, two such (families of) characteristic curves are defined such that:

$$\frac{\mathrm{d}t}{\mathrm{d}s} = 1$$
,  $\frac{\mathrm{d}x}{\mathrm{d}s} = u \pm c$ .

This (with +) is used for the first and (with -) for the second equation:

$$\left[ \frac{\mathrm{d}t}{\mathrm{d}s} \frac{\partial}{\partial t} + \frac{\mathrm{d}x}{\mathrm{d}s} \frac{\partial}{\partial x} \right] (u \pm 2c) = 0 \quad \xrightarrow{\text{the chain rule}} \quad \left( \frac{\mathrm{d}}{\mathrm{d}s} \left( u \pm 2c \right) = 0 \right).$$

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**General property:**  $u \pm 2c$  is constant along 'positive'/'negative' characteristic curves defined by  $\frac{dx}{dt} = u \pm c$ .

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Within the framework of the theory of finite-amplitude waves on shallow water the following problems can be solved:

- the dam-break flow,
- the formation of a bore,
- the hydraulic jump.

### **Outline**

### 1 Introduction

- The notion of wave
- Basic wave phenomena
- Mathematical description of a traveling wave

#### 2 Water waves

- Surface waves on deep water
- Dispersion and the group velocity
- Capillary waves
- Shallow-water finite-amplitude waves

#### 3 Sound waves

- Introduction
- Acoustic wave equation
- The speed of sound
- Sub- and supersonic flow

### Sound waves: introduction

**Sound waves** propagate due to the **compressibility** of a medium  $(\nabla \cdot u \neq 0)$ . Depending on frequency one can distinguish:

- infrasound waves below 20 Hz,
- acoustic waves from 20 Hz to 20 kHz,
- ultrasound waves above 20 kHz.

**Acoustics** deals with vibrations and waves in compressible continua in the **audible frequency range**, that is, from 20 Hz (16 Hz) to 20 000 Hz.

### Types of waves in compressible continua:

- an inviscid compressible fluid (only) longitudinal waves,
- an infinite **isotropic solid** longitudinal and shear waves,
- an **anisotropic solid** wave propagation is more complex.

### Assumptions:

- Gravitational forces can be neglected so that the equilibrium (undisturbed-state) pressure and density take on uniform values,  $p_0$  and  $\varrho_0$ , throughout the fluid.
- Dissipative effects, that is viscosity and heat conduction, are neglected.
- The medium (fluid) is homogeneous, isotropic, and perfectly elastic.

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### Small-amplitudes assumption

Particle velocity is small, and there are only very small perturbations (fluctuations) to the equilibrium pressure and density:

$${\it u} - {\rm small}\,, \qquad p = p_0 + \tilde{p} \quad (\tilde{p} - {\rm small})\,, \qquad \varrho = \varrho_0 + \tilde{\varrho} \quad (\tilde{\varrho} - {\rm small})\,.$$

The pressure fluctuations field  $\tilde{p}$  is called the **acoustic pressure**.

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Momentum equation (Euler's equation):

$$\varrho \left( \, rac{\partial {m u}}{\partial t} + {m u} \cdot 
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Notice that  $\nabla p = \nabla (p_0 + \tilde{p}) = \nabla \tilde{p}$ .

Continuity equation:

$$\frac{\partial \varrho}{\partial t} + \nabla \cdot (\varrho \, \pmb{u}) = 0 \quad \xrightarrow{\text{linearization}} \quad \frac{\partial \tilde{\varrho}}{\partial t} + \varrho_0 \, \nabla \cdot \pmb{u} = 0 \, .$$

Momentum equation (Euler's equation):

$$\varrho\bigg(\,\frac{\partial \textbf{\textit{u}}}{\partial t} + \textbf{\textit{u}}\cdot\nabla\textbf{\textit{u}}\,\bigg) = -\nabla p \quad \xrightarrow{\text{linearization}} \quad \varrho_0\;\frac{\partial \textbf{\textit{u}}}{\partial t} = -\nabla p\,.$$

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Using divergence operation for the linearized momentum equation and time-differentiation for the linearized continuity equation yields:

$$\frac{\partial^2 \tilde{\varrho}}{\partial t^2} - \triangle p = 0.$$

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$$p = p(\tilde{\varrho}) \quad \rightarrow \quad \frac{\partial p}{\partial t} = \frac{\partial p}{\partial \tilde{\varrho}} \; \frac{\partial \tilde{\varrho}}{\partial t} \quad \rightarrow \quad \frac{\partial^2 \tilde{\varrho}}{\partial t^2} = \frac{1}{c_0^2} \; \frac{\partial^2 p}{\partial t^2} \quad \text{where } c_0^2 = \frac{\partial p}{\partial \tilde{\varrho}} \; .$$

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### Wave equation for the pressure field

$$\boxed{ \frac{1}{c_0^2} \; \frac{\partial^2 p}{\partial t^2} - \triangle p = 0 } \quad \text{where} \quad c_0 = \sqrt{\frac{\partial p}{\partial \tilde{\varrho}}}$$

is the **acoustic wave velocity** (or the **speed of sound**). Notice that the acoustic pressure  $\tilde{p}$  can be used here instead of p. Moreover, the wave equation for the density-fluctuation field  $\tilde{\varrho}$  (or for the compression field  $\tilde{\varrho}/\varrho_0$ ), for the velocity potential  $\phi$ , and for the velocity field  $\boldsymbol{u}$  can be derived analogously.

# The speed of sound

**Inviscid isotropic elastic liquid.** The pressure in an inviscid liquid depends on the volume dilatation  ${\rm tr}\, \varepsilon$ :

$$p = -K \operatorname{tr} \boldsymbol{\varepsilon},$$

where K is the bulk modulus. Now,

$$\frac{\partial p}{\partial t} = -K \operatorname{tr} \frac{\partial \boldsymbol{\varepsilon}}{\partial t} = -K \nabla \cdot \boldsymbol{u} \qquad \frac{\nabla \cdot \boldsymbol{u} = -\frac{1}{\varrho_0} \frac{\partial \varrho}{\partial t}}{\text{Lin. Cont. Eq.}} \qquad \frac{\partial p}{\partial t} = \frac{K}{\varrho_0} \frac{\partial \tilde{\varrho}}{\partial t}$$

which means that the speed of sound  $c_0 = \sqrt{\partial p/\partial \tilde{\varrho}}$  is given by the well-known formula:

$$c_0 = \sqrt{\frac{K}{\varrho_0}}$$

# The speed of sound

**Inviscid isotropic elastic liquid.** The speed of sound is given by the well-known formula:

$$c_0 = \sqrt{\frac{K}{\varrho_0}}$$

**Perfect gas.** The determination of speed of sound in a perfect gas is complicated and requires the use of thermodynamic considerations. The final result is

$$c_0 = \sqrt{\gamma \frac{p_0}{\varrho_0}} = \sqrt{\gamma R T_0},$$

where  $\gamma$  denotes the ratio of specific heats ( $\gamma=1.4$  for air), R is the universal gas constant, and  $T_0$  is the (isothermal) temperature.

▶ For air at 20°C and normal atmospheric pressure:  $c_0 = 343 \, \frac{\text{m}}{\text{s}}$ .

A steady, unseparated, **compressible flow** past a thin airfoil may be written in the from

$$u = U + \frac{\partial \phi}{\partial x}$$
,  $v = \frac{\partial \phi}{\partial y}$ ,

where the **velocity potential**  $\phi$  for the small disturbance to the uniform flow U satisfies

$$(1-M^2) \; rac{\partial^2 \phi}{\partial x^2} + rac{\partial^2 \phi}{\partial y^2} = 0 \,, \qquad ext{where} \qquad \overline{M = rac{U}{c_0}}$$

is the **Mach number** defined as the ratio of the speed of free stream to the speed of sound.

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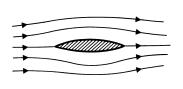
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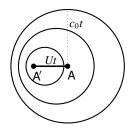
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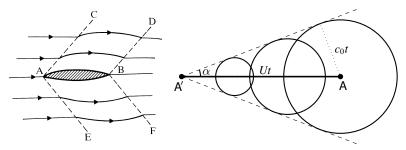
- ▶ If  $M^2 \ll 1$  that gives the Laplace equation which is the result that arises for incompressible theory (i.e., using  $\nabla \cdot \boldsymbol{u} = 0$ ).
- Otherwise, three cases can be distinguished:
  - 1 M < 1 the subsonic flow
  - 2 M > 1 the supersonic flow
  - 3  $M \approx 1$  the sound barrier

- 1 M < 1 the subsonic flow:
  - there is some disturbance to the oncoming flow at all distances from the wing (even though it is very small when the distance is large);
  - the **drag is zero** (inviscid theory) and the lift =  $\frac{\text{lift}_{\text{incompressible}}}{\sqrt{1-M^2}}$





- 1 M < 1 the subsonic flow
- **2** M > 1 the supersonic flow:
  - there is no disturbance to the oncoming stream except between the **Mach lines** extending from the ends of the airfoil and making the angle  $\alpha = \arcsin\left(\frac{1}{M}\right)$  with the uniform stream;
  - the **drag is not zero** it arises because of the sound wave energy which the wing radiates to infinity between the Mach lines.



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  - there is some disturbance to the oncoming flow at all distances from the wing (even though it is very small when the distance is large);
  - $\blacksquare$  the **drag is zero** (inviscid theory) and the lift  $= \frac{\text{lift}_{\text{incompressible}}}{\sqrt{1-M^2}}$  .
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  - the **drag is not zero** it arises because of the sound wave energy which the wing radiates to infinity between the Mach lines.
- 3  $M \approx 1$  the sound barrier:
  - sub- and supersonic theory is not valid;
  - nonetheless, it indicates that the wing is subject to a destructive effect of exceptionally large aerodynamic forces.